

Indirect WIMP detection with neutrinos in Hyper-K

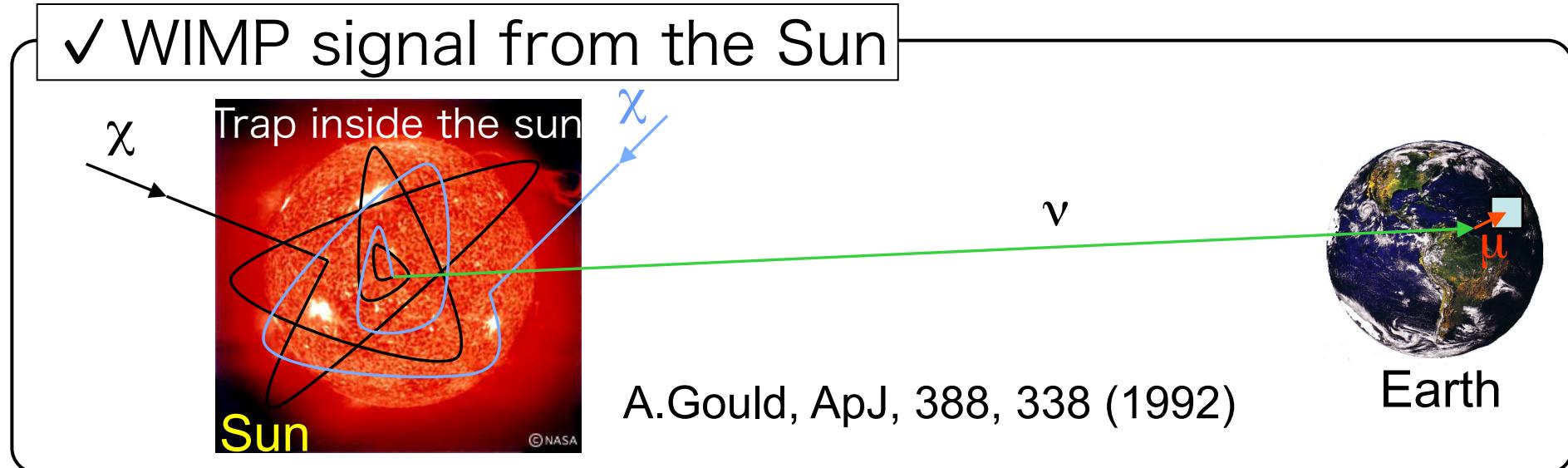
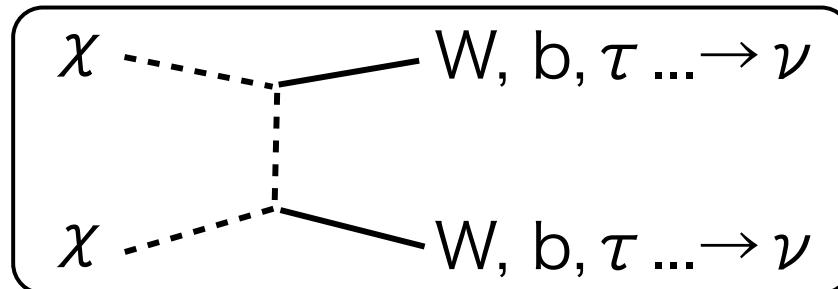
Yusuke Koshio

for Hyper-K astrophysics working group
Kamioka observatory, ICRR, Univ. of Tokyo

Cosmic Frontier Workshop

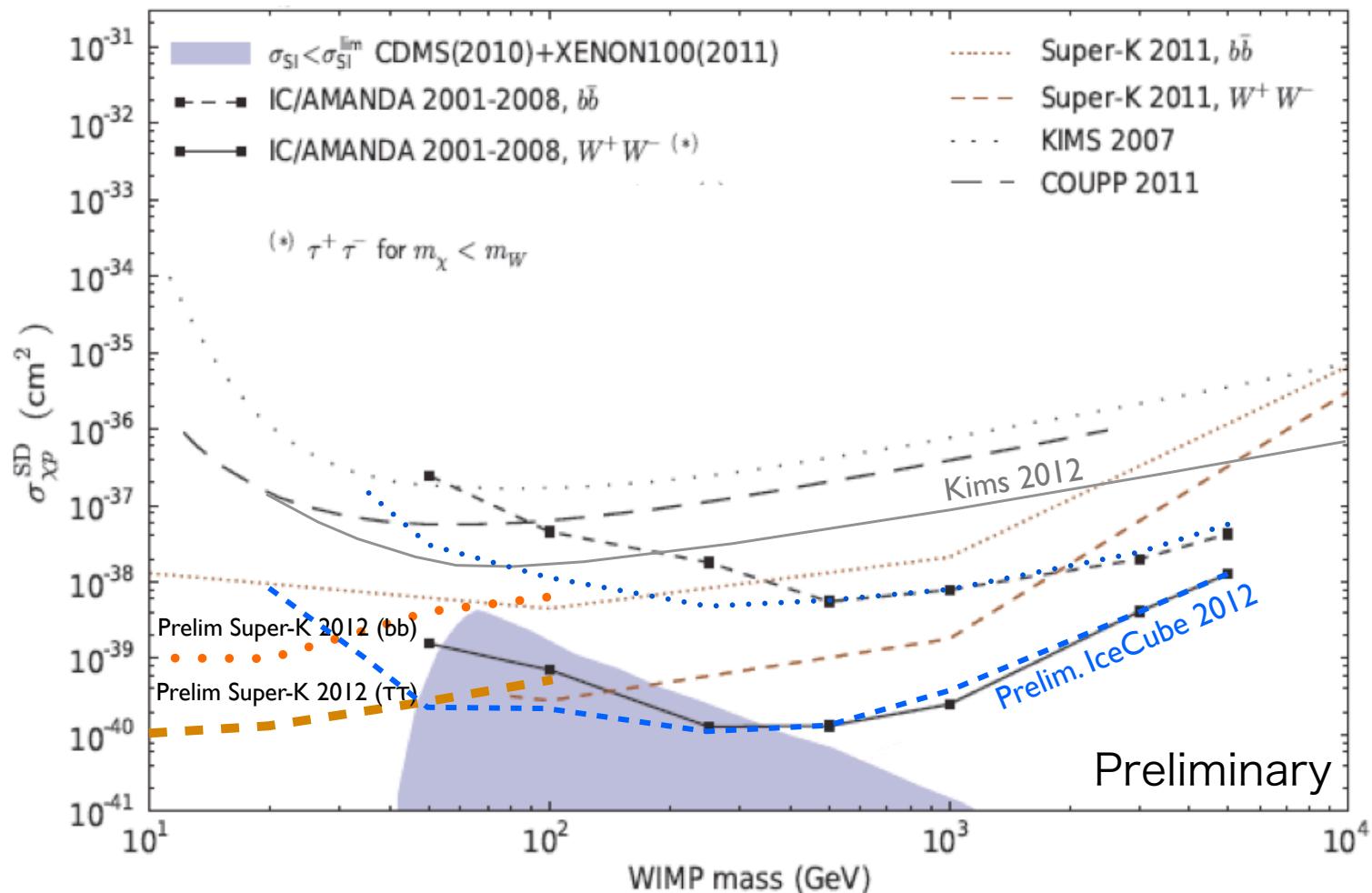
SLAC - March 7, 2013

Neutrinos generated by WIMP annihilation



✓ Annihilation in Galactic Halo, etc..

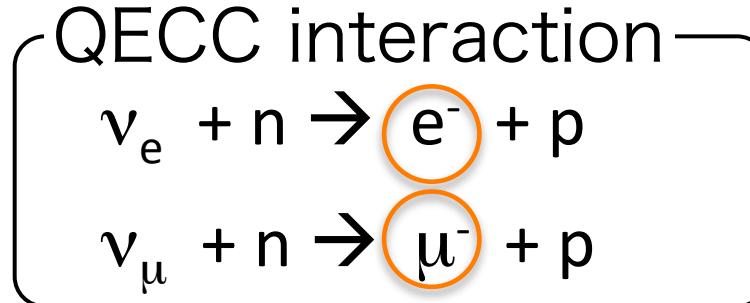
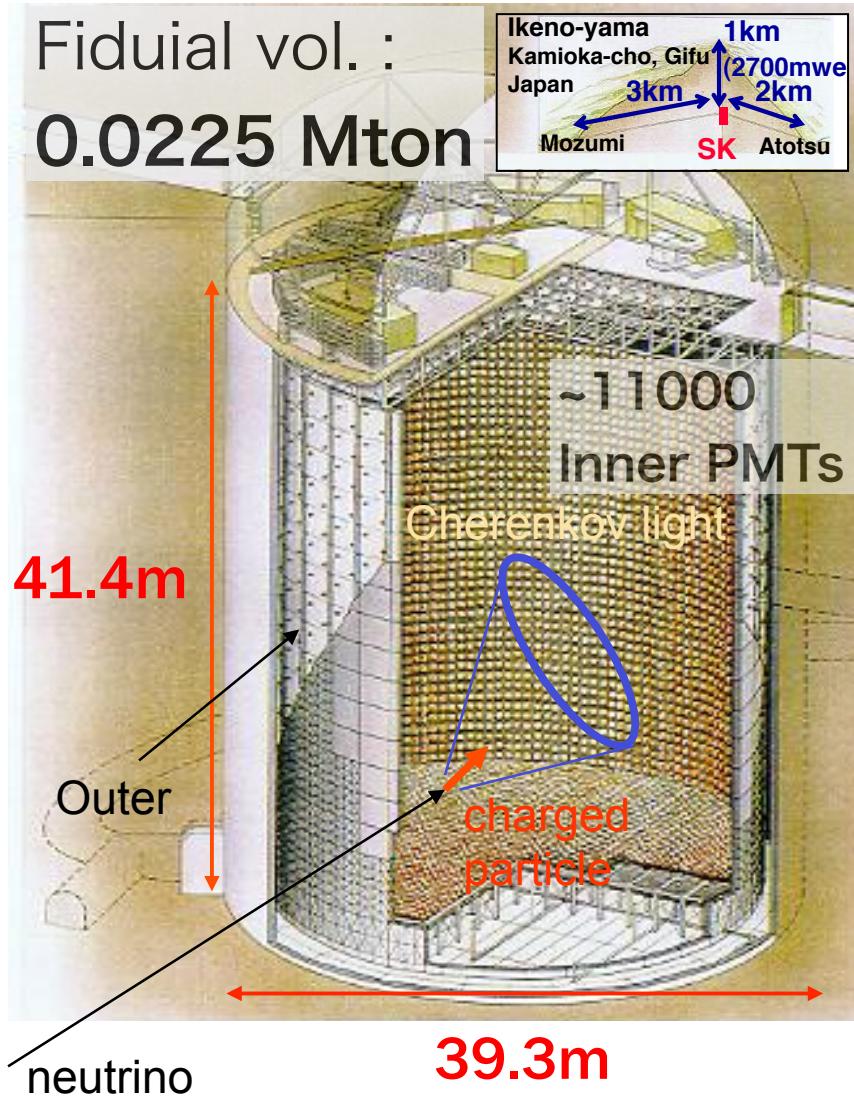
Recent results



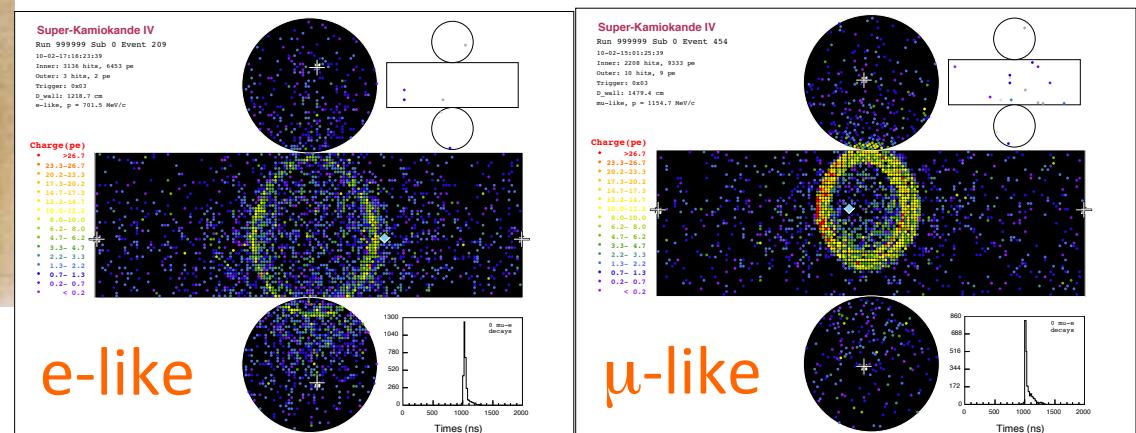
T. Tanaka et al. *Astrophys. J.* **742**, 78 (2011)
 R. Abbasi et al. *Phys. Rev. D* **85**, 042002 (2012)

Preliminary IceCube/DeepCore Limit IDM 2012 / arXiv:1212.4097
 Preliminary Super-K Limit Neutrino 2012 / arXiv:1210.4161

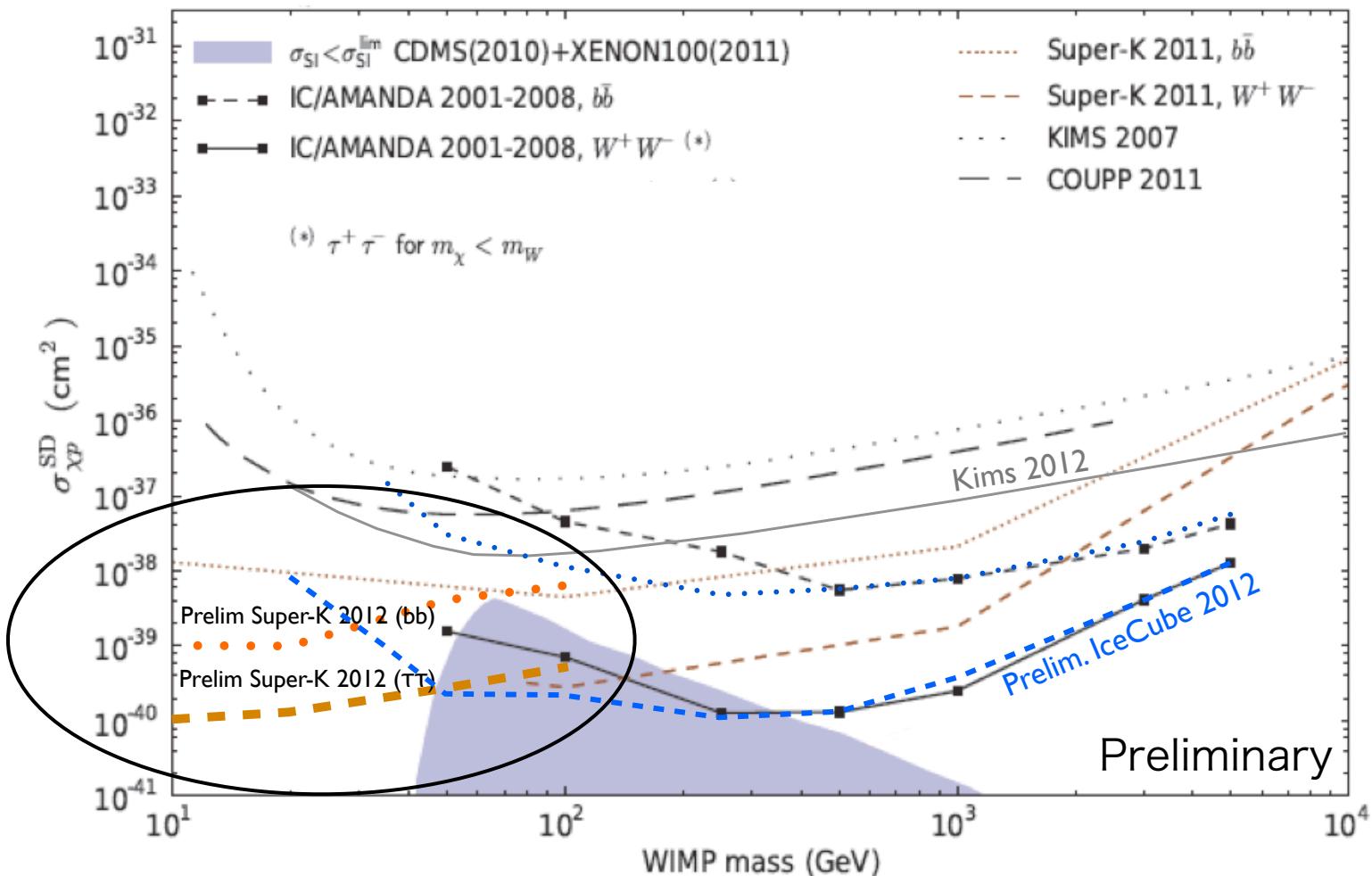
Observation in SuperK



- ✓ Good Particle ID
- ✓ Conserve the neutrino direction
 - Important to search for the signal from the Sun
- ✓ Good energy reconstruction



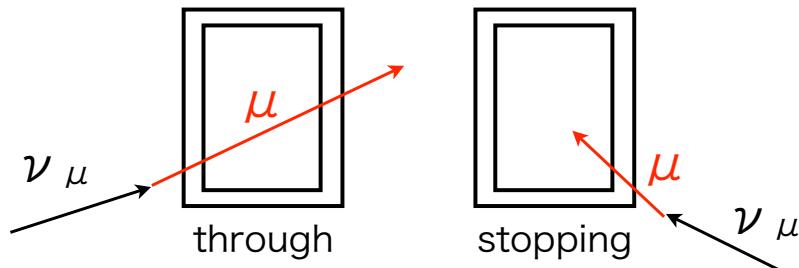
Recent results



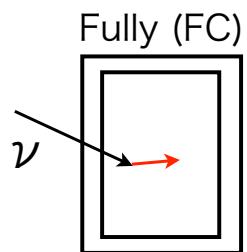
Observation in SuperK

Event category

✓ Upward going muons (Upmu)

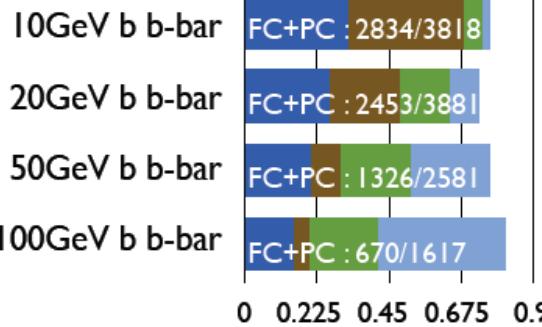
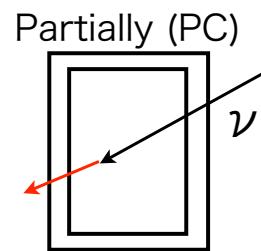


✓ Contained events

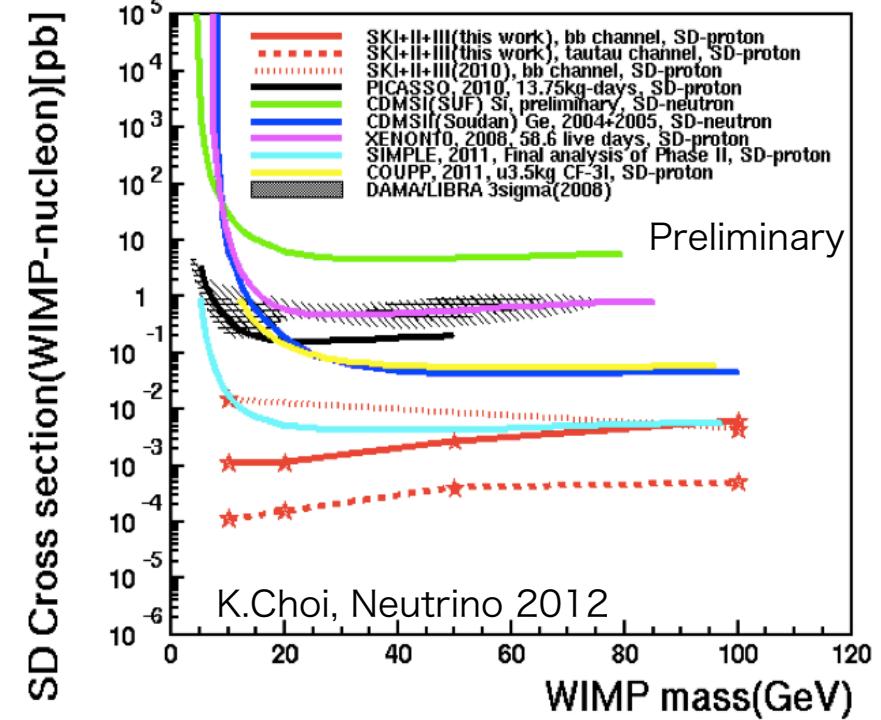
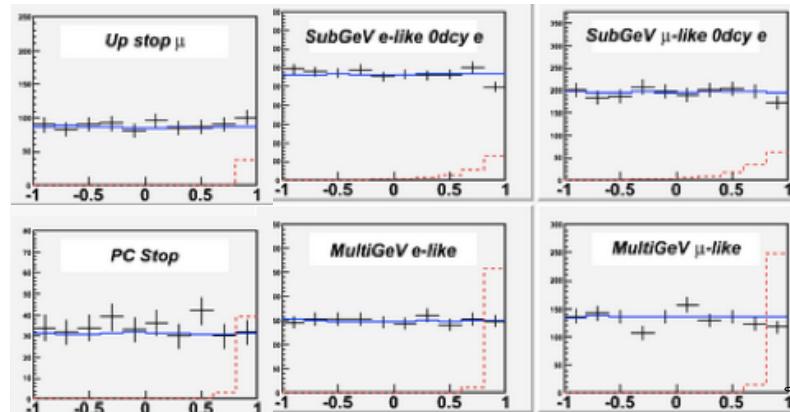


sensitive to low-mass

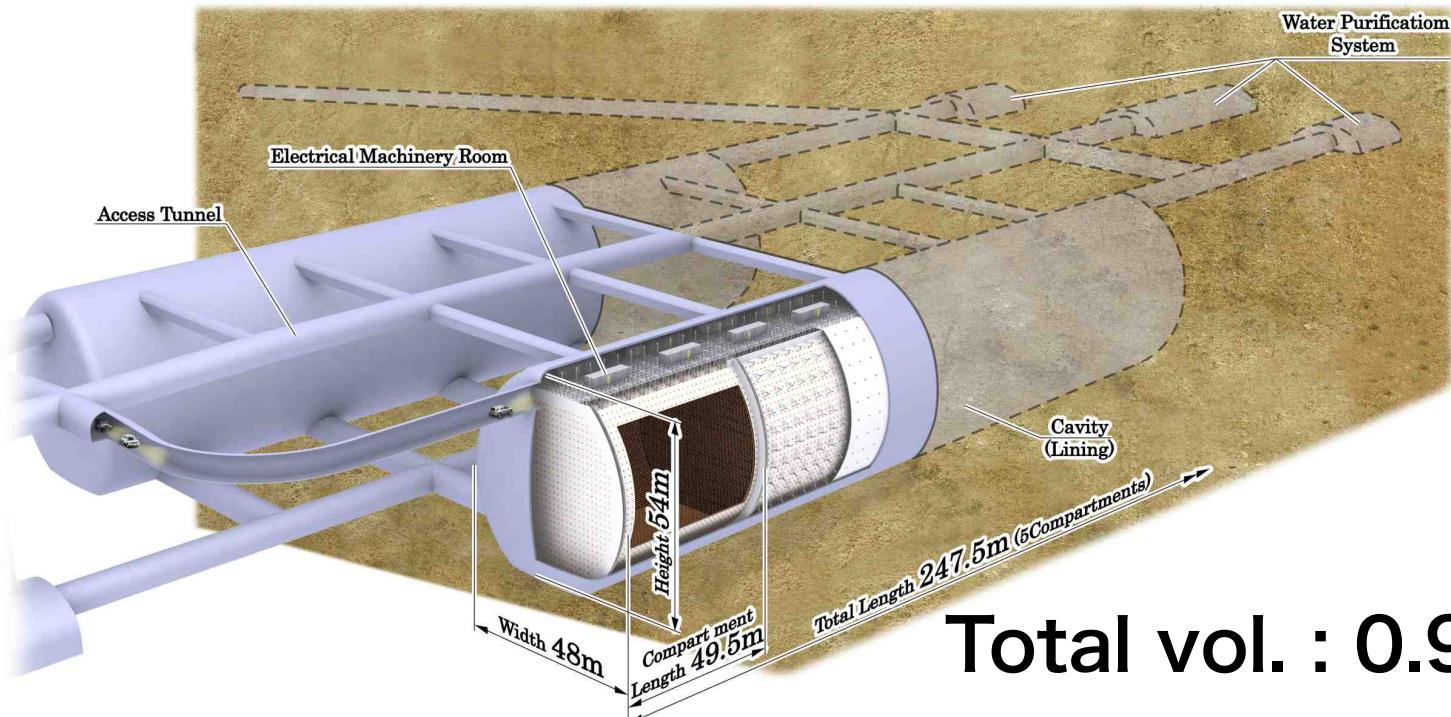
- FC I-ring e-like
- FC I-ring m-like
- PC
- Upmu



No significant excess from the Sun



Hyper-Kamiokande

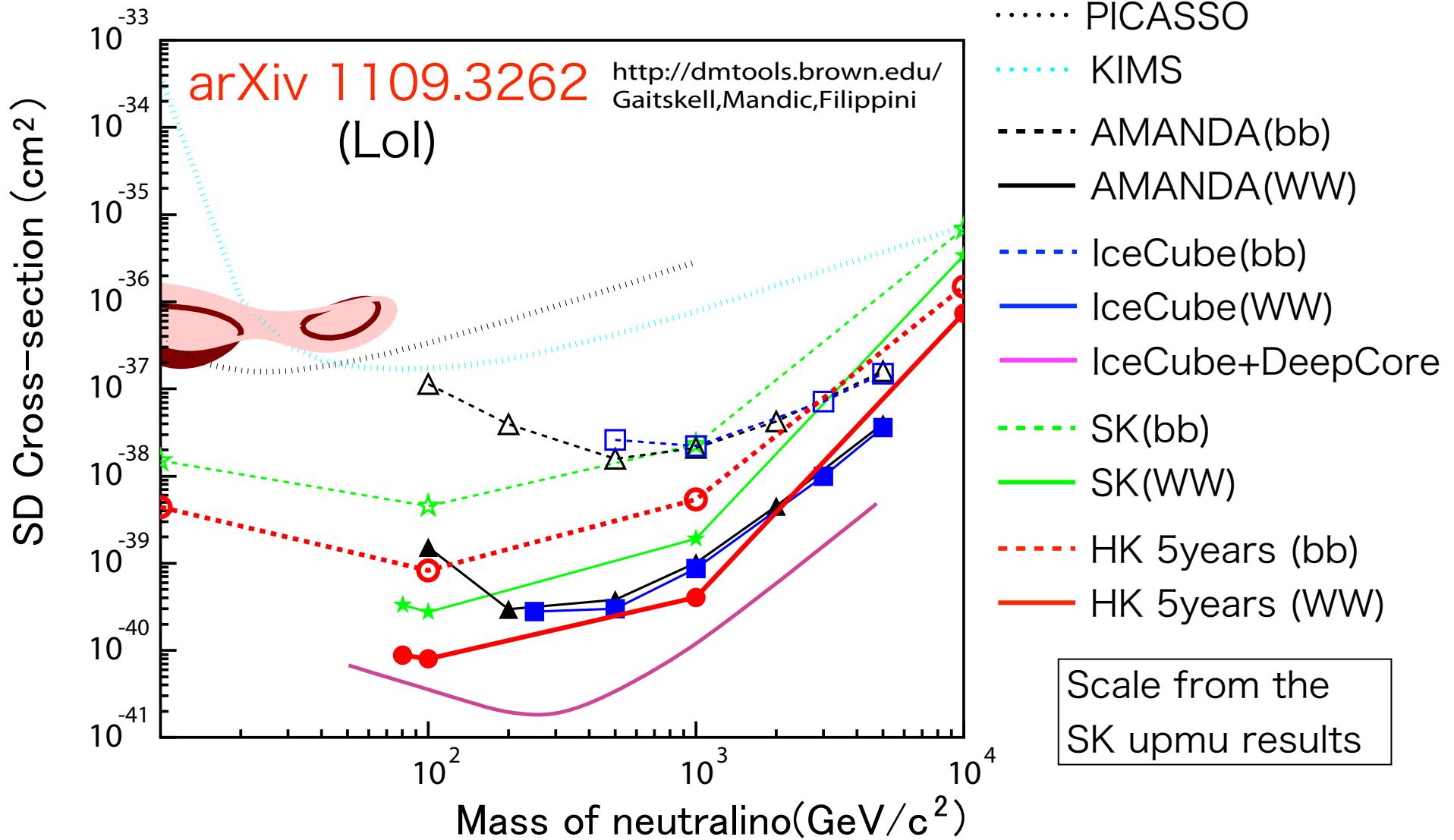


Lol : arXiv 1109.3262

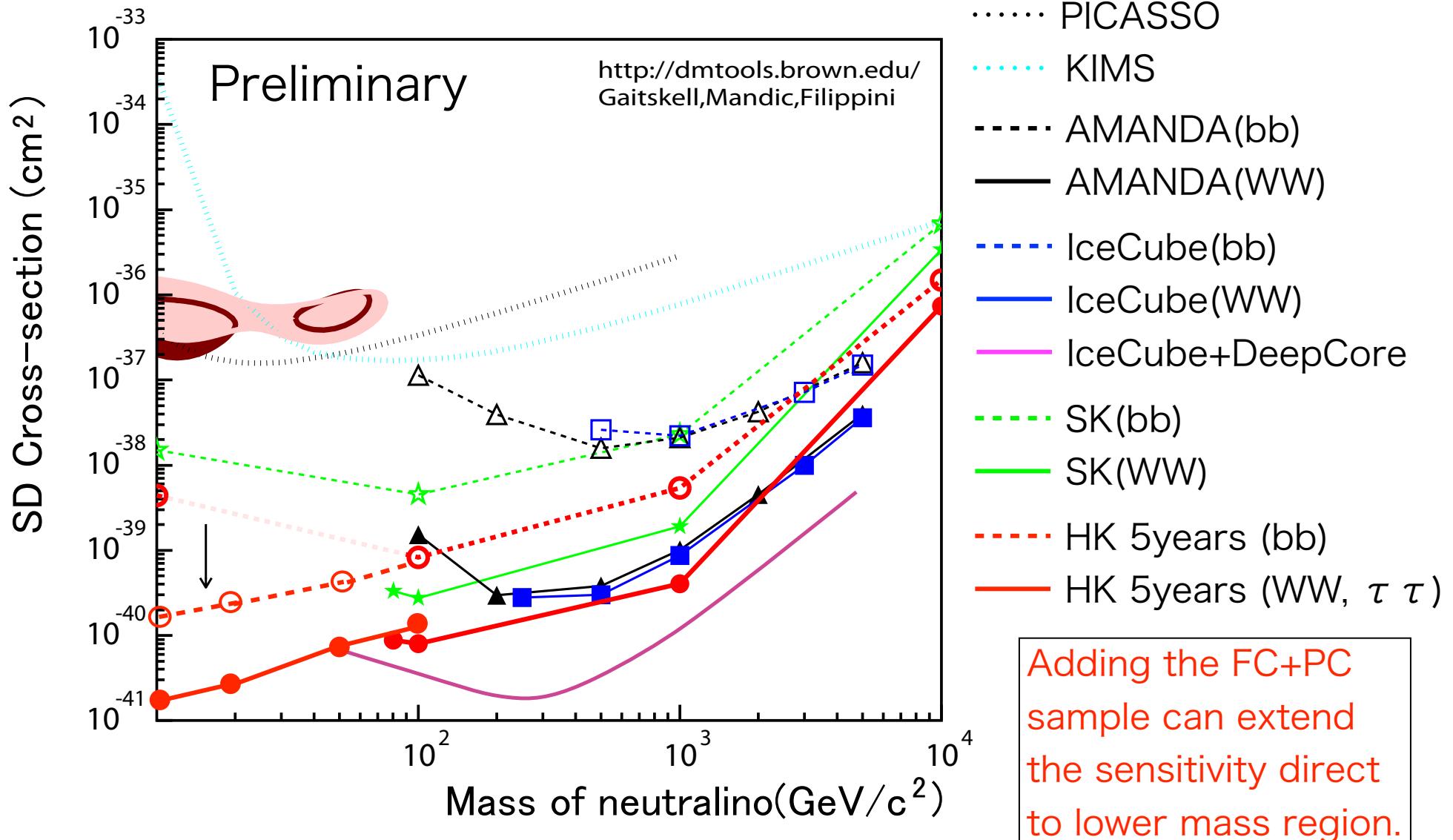
Total vol. : 0.99 Mton
Fiducial vol. : 0.56 Mton
Photo-sensors : 99,000 PMTs

25 times larger than Super-K

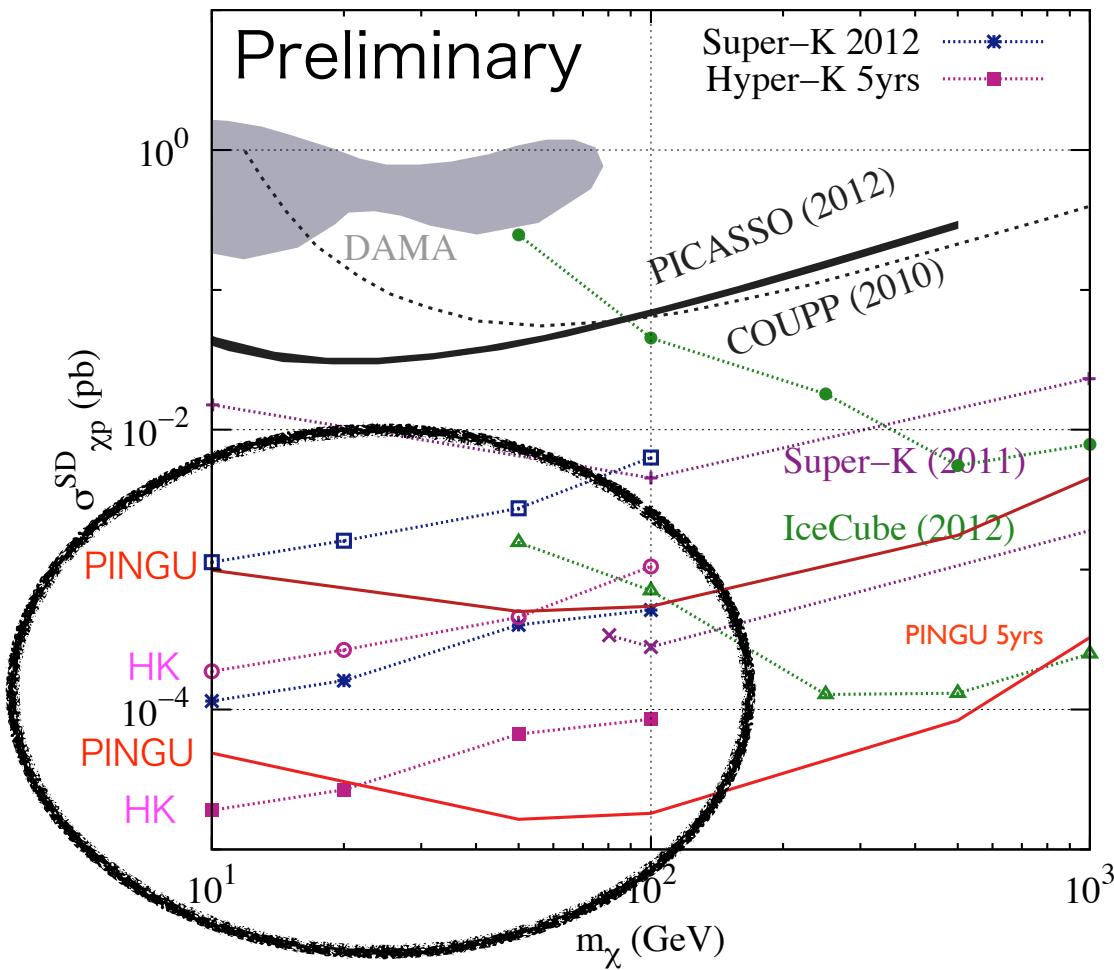
Hyper-K sensitivity



Hyper-K sensitivity



Hyper-K sensitivity



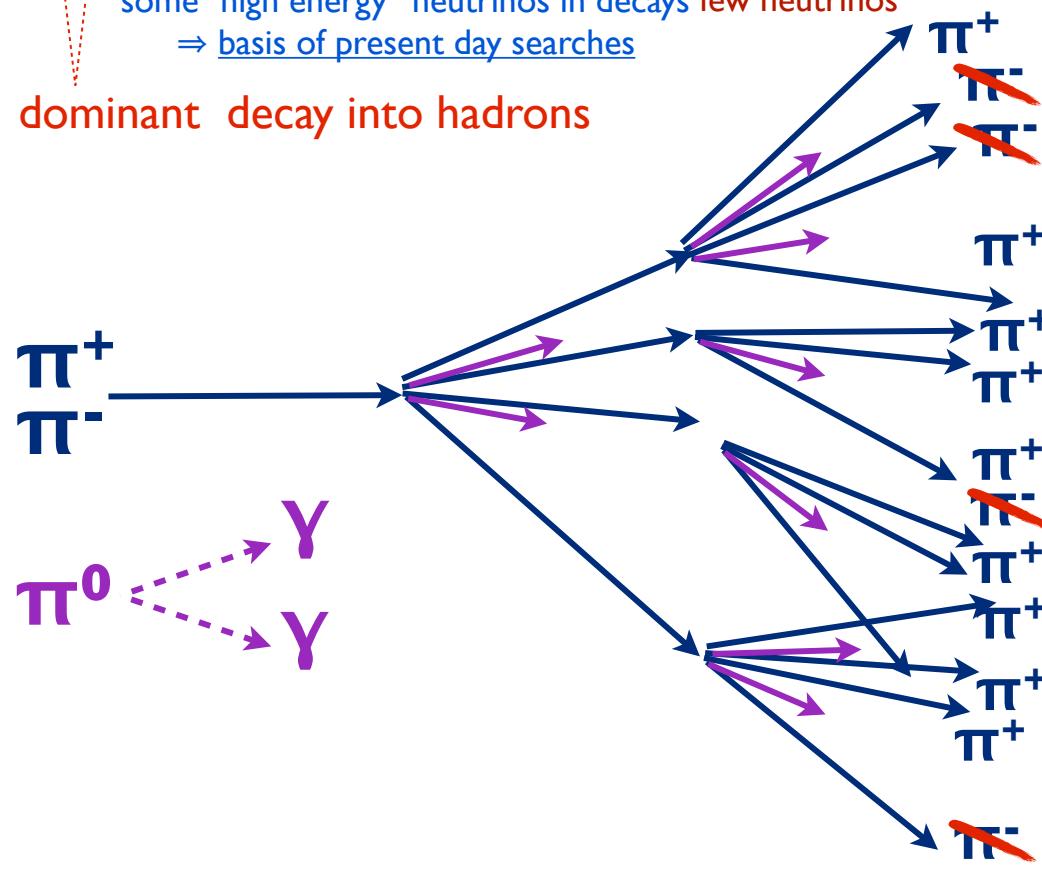
- Scale of achieved Super-K results (FC+PC+upmu) to Hyper-K.
- Hyper-K is sensitive to the interesting WIMP mass region suggested by the direct search experiments.
- Hyper-K compares favorable to other indirect searches for WIMP masses below ~ 20 - 50 GeV.

New detection channel

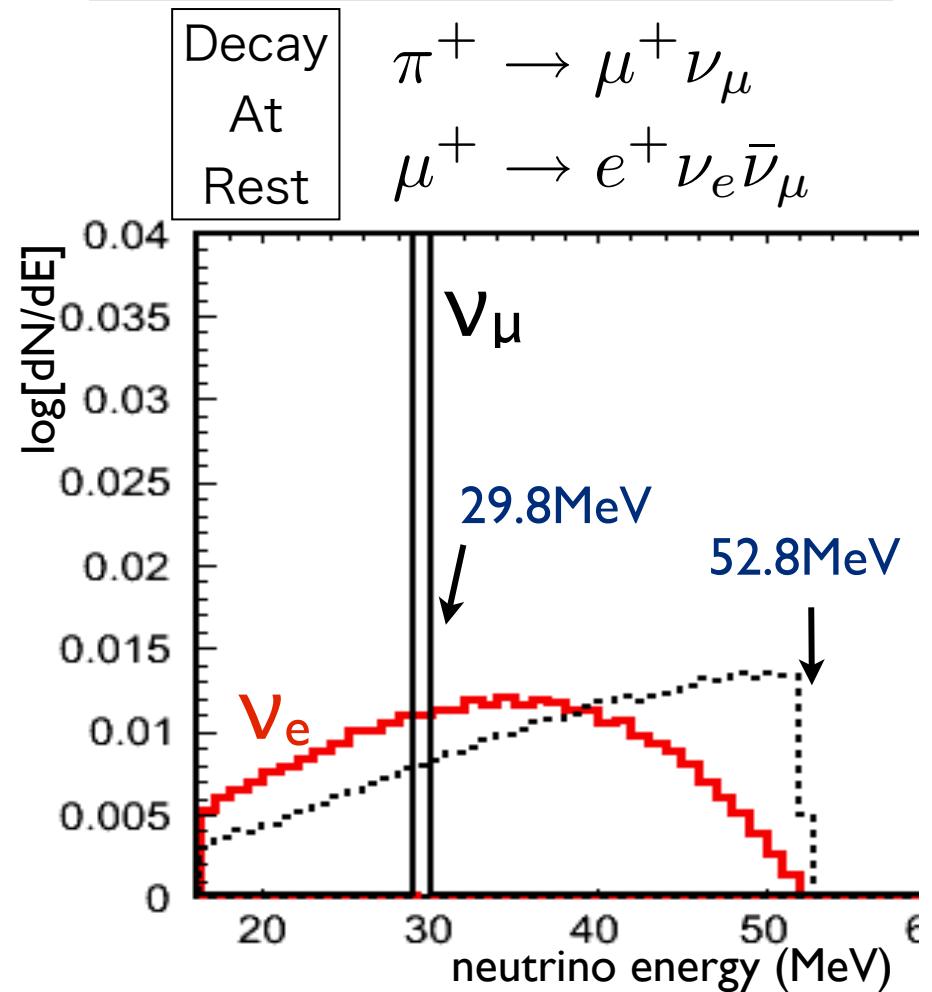
Possible annihilation channels:
 $qq, gg, cc, ss, bb, tt, W^+W^-, ZZ, \tau^+\tau^-, \mu^+\mu^-, \nu\nu, e^+e^-, \gamma\gamma$

some “high energy” neutrinos in decays few neutrinos
 \Rightarrow basis of present day searches

dominant decay into hadrons

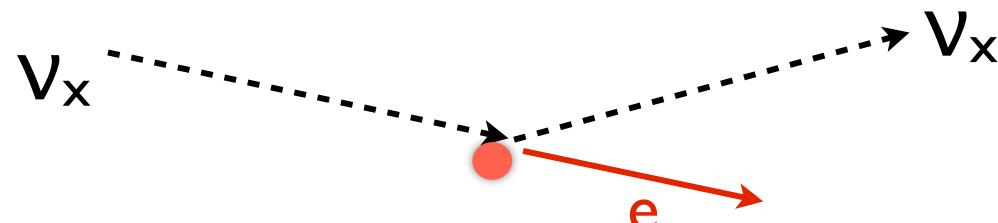


Charged pions at rest producing neutrinos up to $E=52.8\text{MeV}$

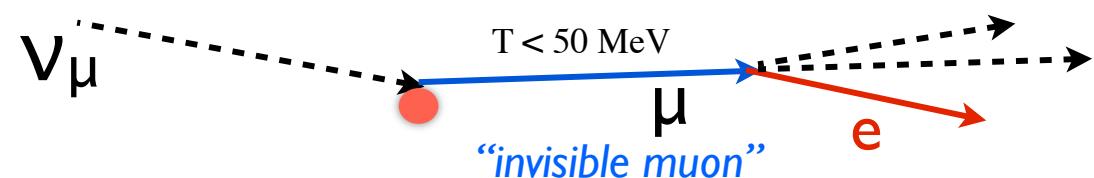


Backgrounds

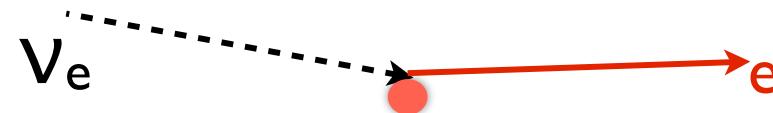
NC Elastic
“reactor + solar”



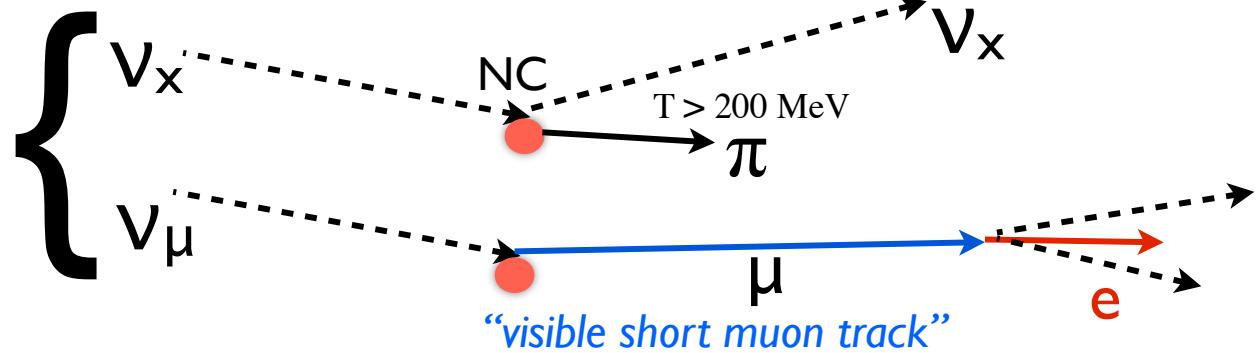
Decay electron
“atm. muon neutrinos”



ν_e CC
“atm. electron neutrinos”

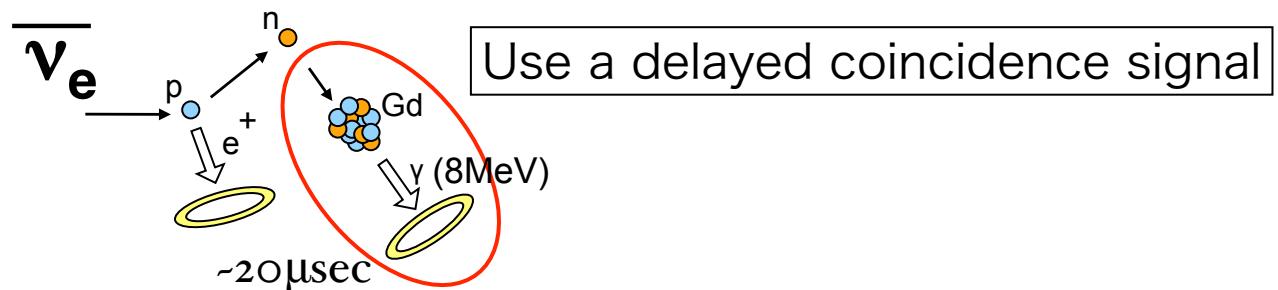


μ/π
“ μ/π production from atm. neutrinos”



Gadolinium in water

- Decay electron events are the dominant background
- Identifying neutrinos of the inverse beta decay reaction can provide a way to discriminate against this background

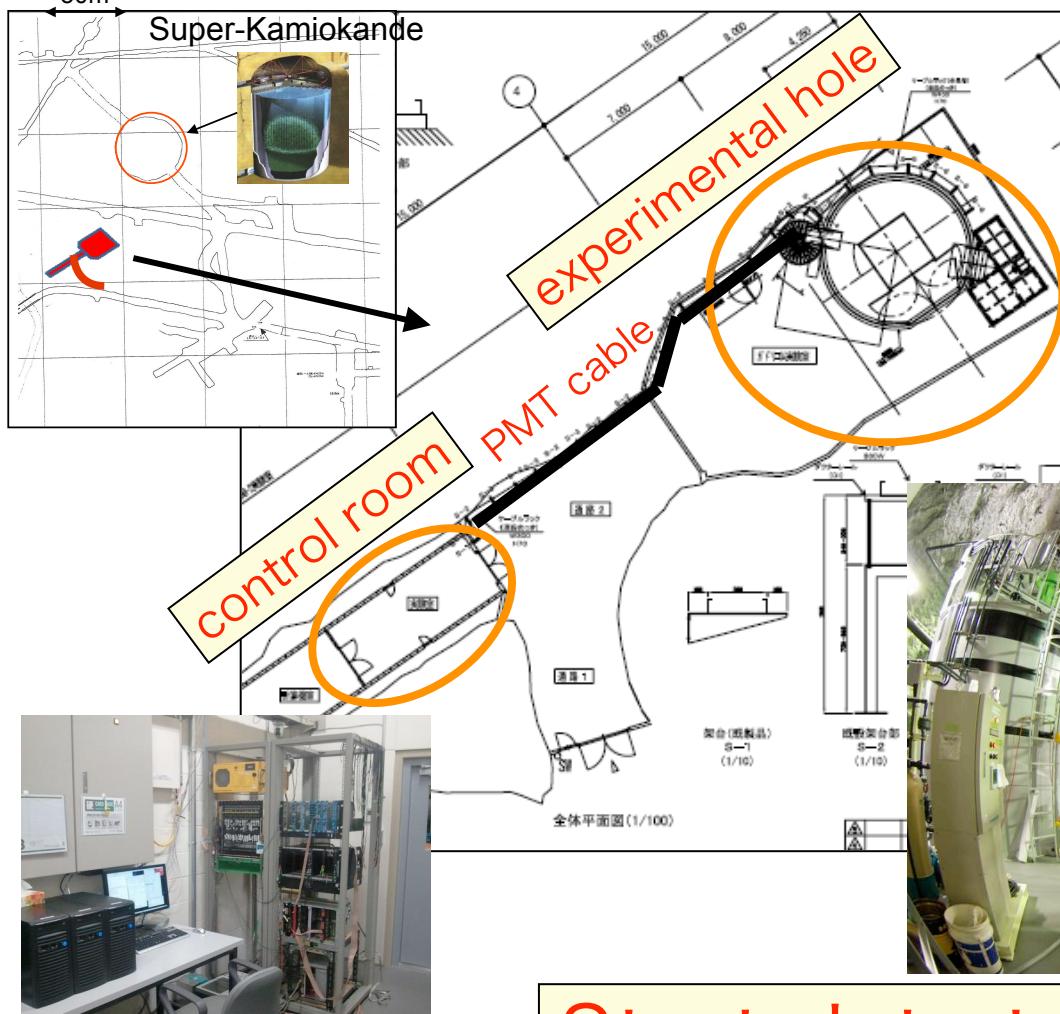


- Proposal : add Gd to Super-K for SuperNova relic neutrino search. (Beacom & Vagins : Phys.Rev.Lett. 93(2004)171101)
 - 0.2% Gd solution would yield >90% neutron captured.

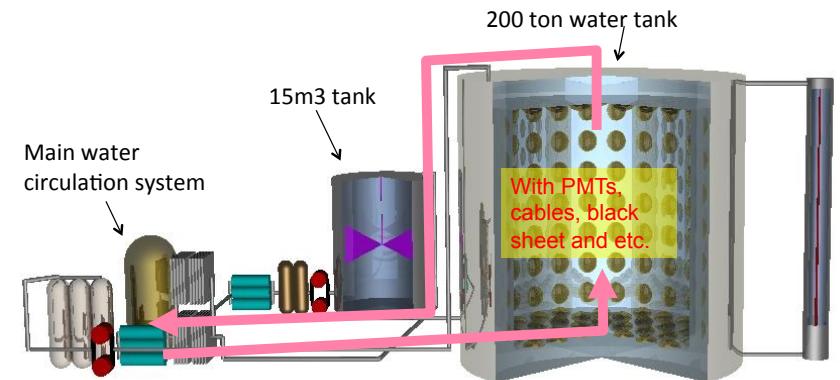
EGADS

Evaluating Gadolinium's Action on Detector Systems

50m Kamioka underground observatory

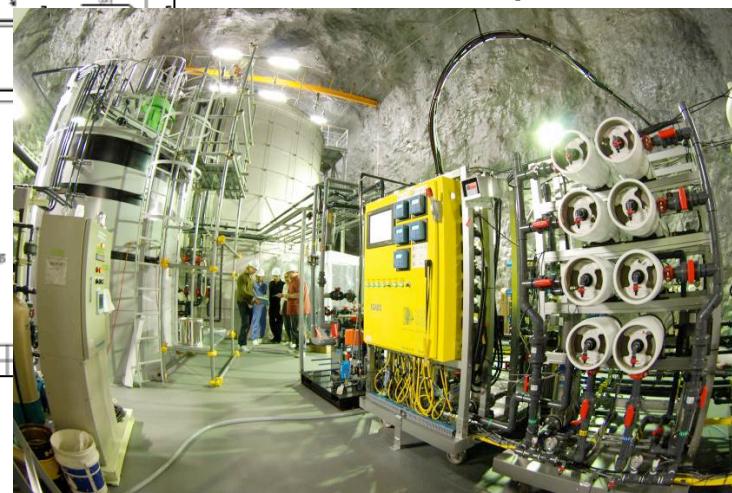


Electronics and computer



R&D for Gd in SuperK
(GADZOOKS!)

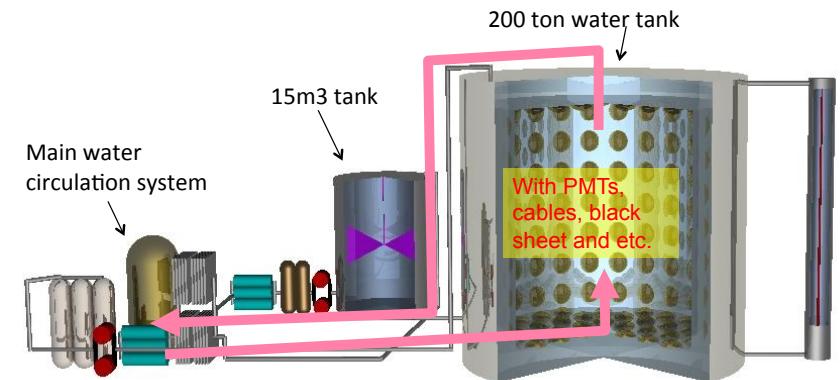
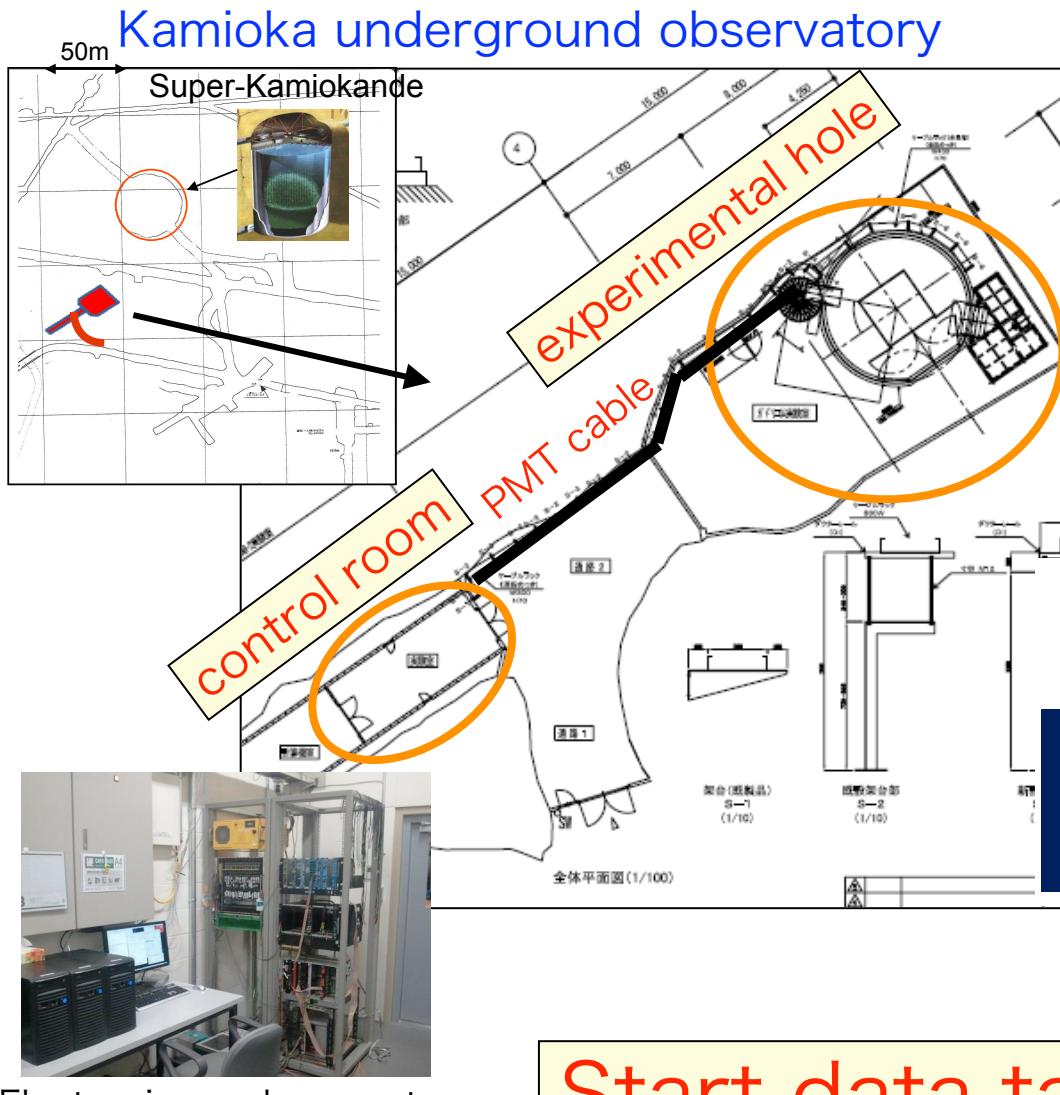
- 200 ton tank with 232 PMTs
- Establish purification method



Start data taking in this Summer

EGADS

Evaluating Gadolinium's Action on Detector Systems



R&D for Hyper-K (Gd option)

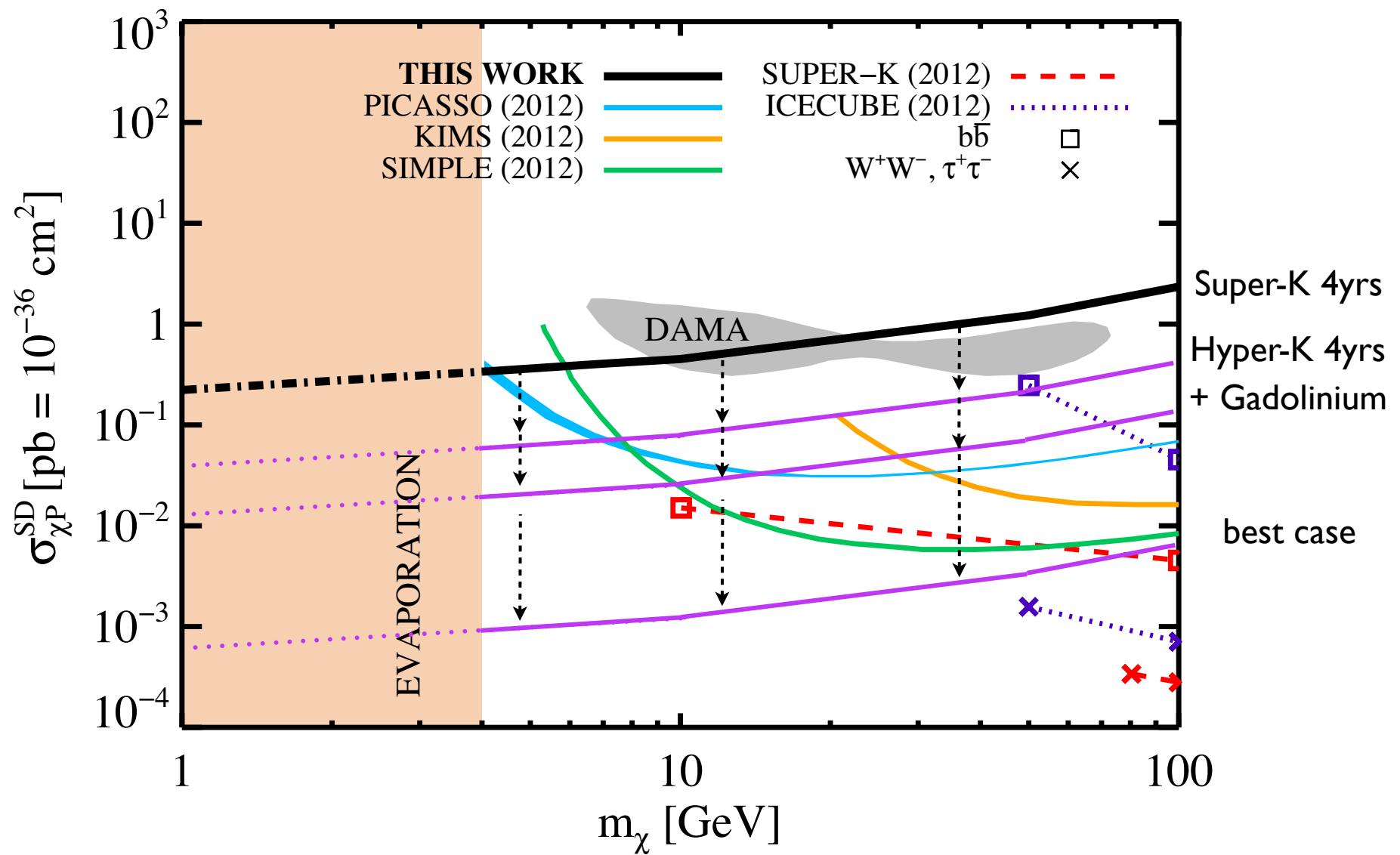
- 200 ton tank with 8 HPDs
- Establish purification method



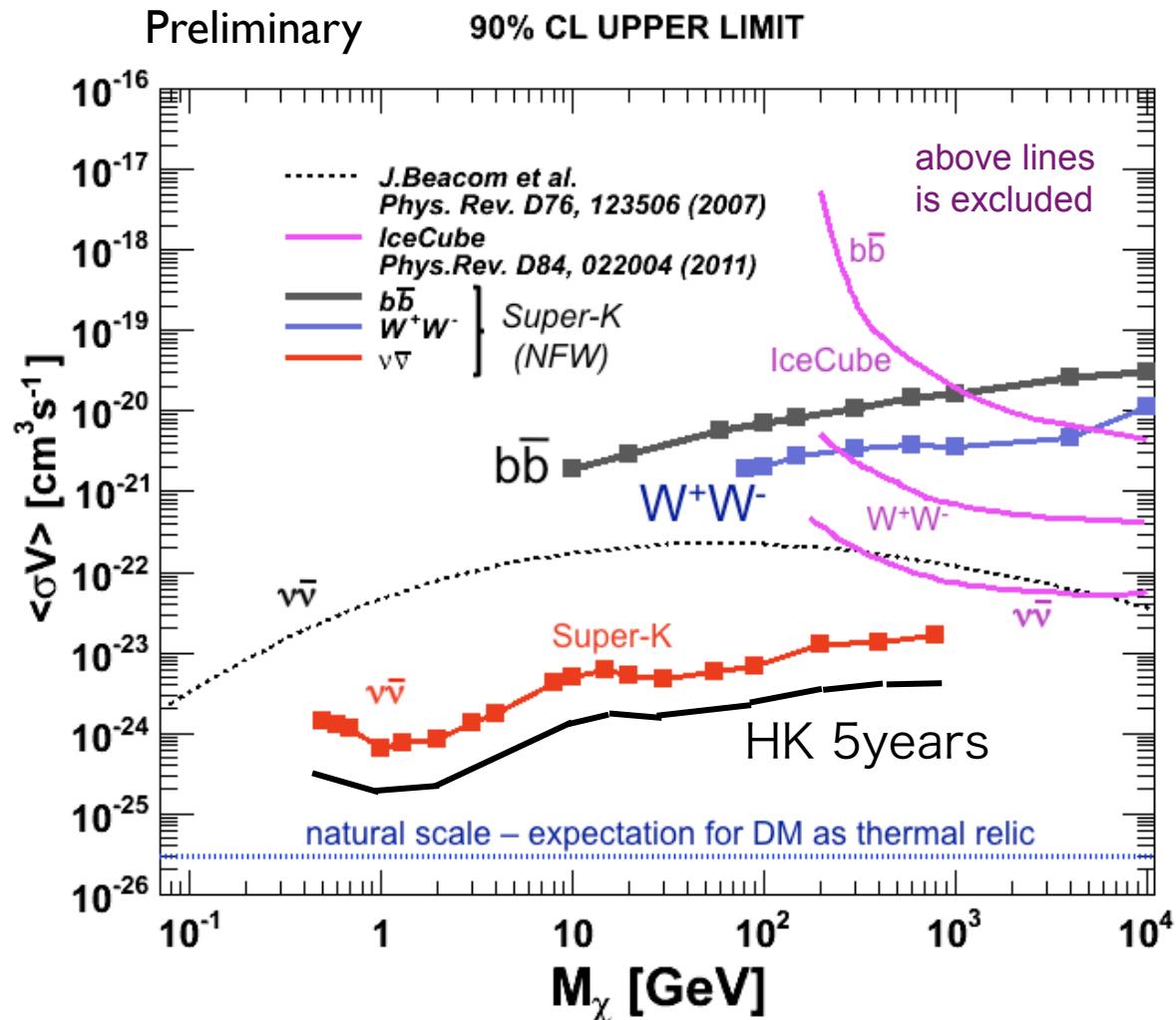
M.Yokoyama, this afternoon session

Start data taking in this Summer

Hyper-K sensitivity



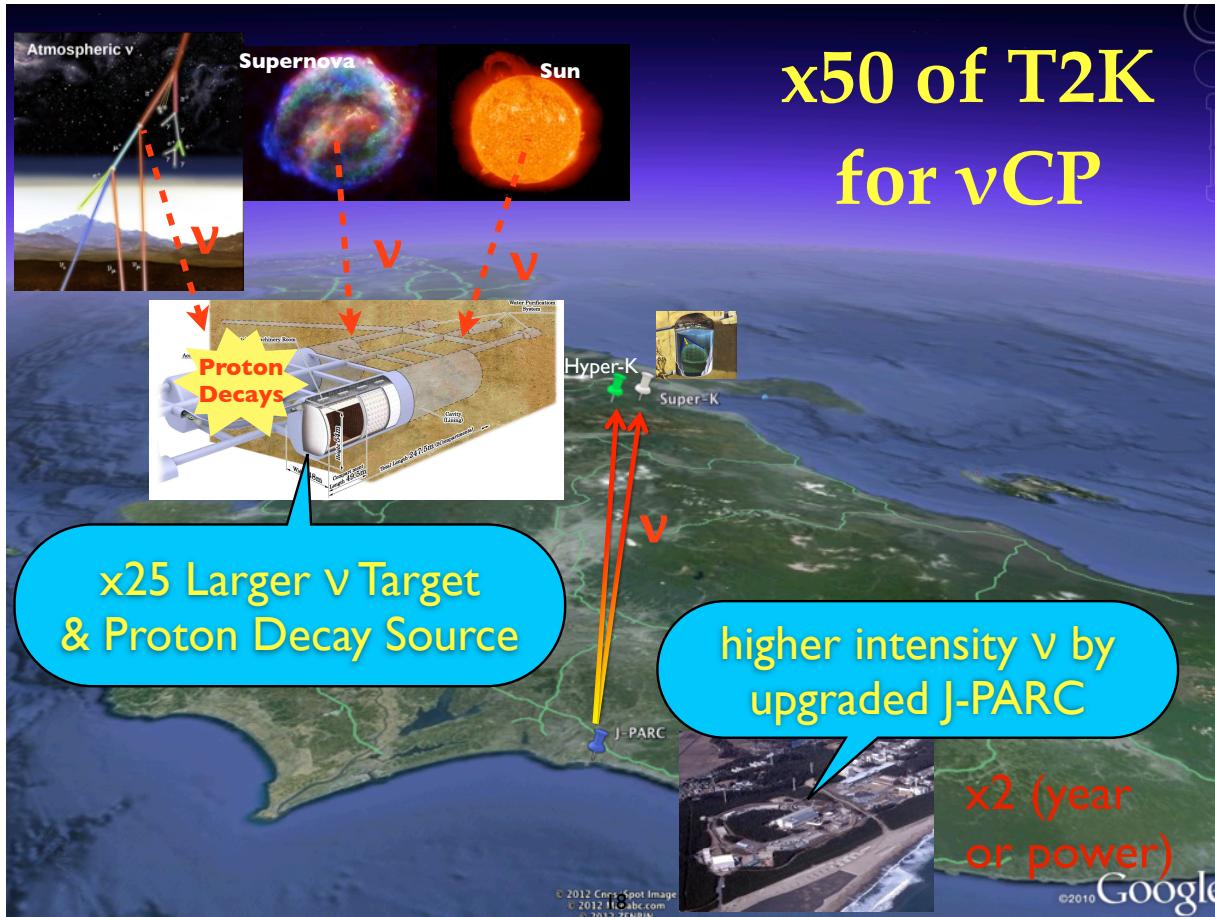
Galactic search



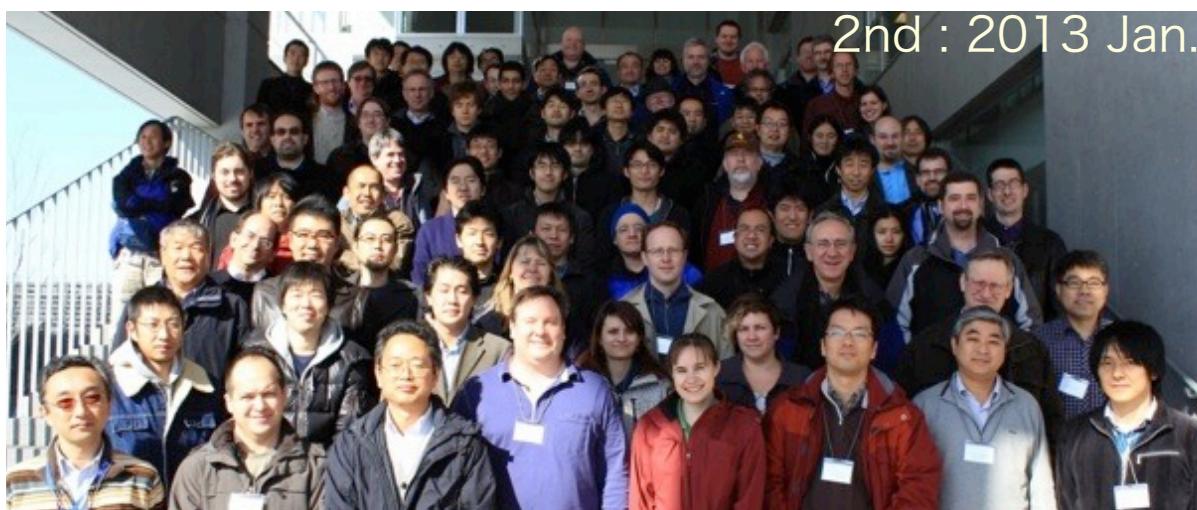
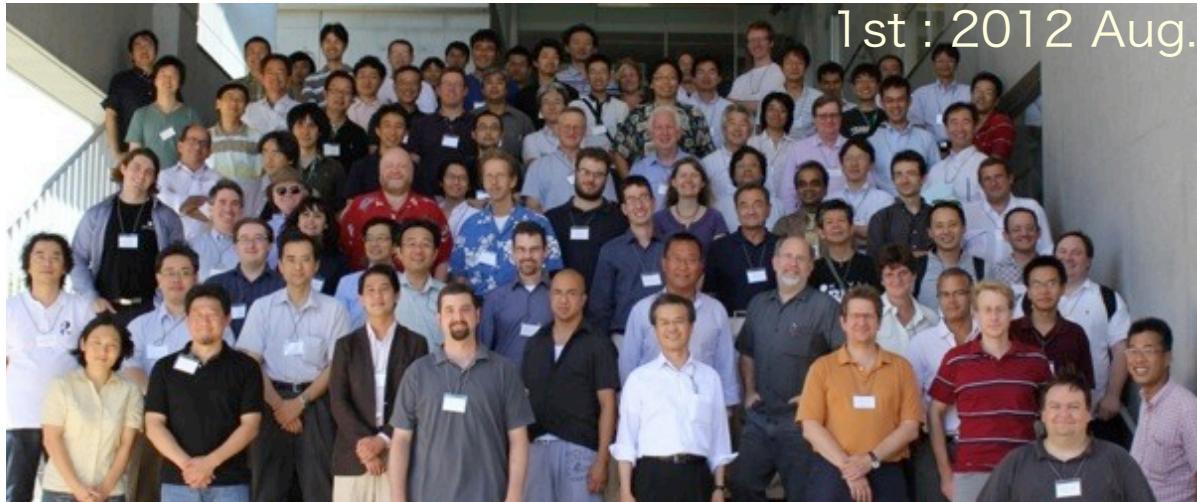
Hyper-Kamiokande

multi-purpose detector

Lol : arXiv 1109.3202



International HK meeting



Start up an international working group (~100 researchers)

< List of R&D >

- Physics potential
 - Cavity and Tank
 - Water system
 - Photo-sensor development
 - DAQ and electronics
 - software
 - calibration
- etc.

Conclusion

- Water Cherenkov detector has a unique potential for indirect WIMP search.
 - Established by Super-K.
 - Good flavor tagging, energy and direction reconstruction.
- Hyper-K will be most important to cover the WIMP mass region less than 50GeV.
- New low-energy neutrinos from the Sun could offer unique prospects for Hyper-K.